EbE Vertexing for Mixing

Alex

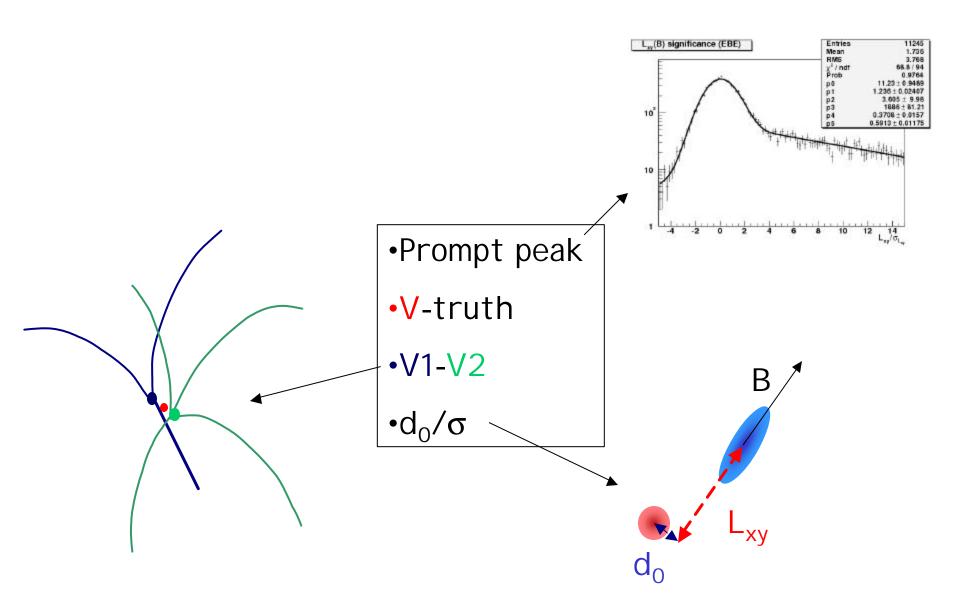


Status

- Increased sample's statistics
 - ✓ Full ~350 pb⁻¹
 - \checkmark D⁰π, D⁺π, ψ K⁺, ψ K^{*}, ψ' \rightarrow μμππ
- Primary Vertex:
 - ✓ SF robustly sitting around 1.38
 - ✓ Dependencies (Z,Pt,Si hits,Ntracks)
 - ✓ Effect of hourglass
 - \checkmark G3X (no time to show today, but does not seem relevant with current statistics)
 - ✓ Systematics
 - ✓ Comparison of pulls and extraction of a common value:
 - D₀ vs L_{xy}
- Secondary Vertex:
 - ✓ Dependencies $(Z,\phi,\eta, Pt,L_{xy},ct,\Delta\phi,\Delta R,I solation,Si hits)$
 - Extraction of a common scale factor with systematics?

Primary Vertex

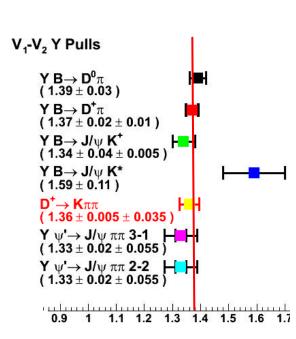
The tools

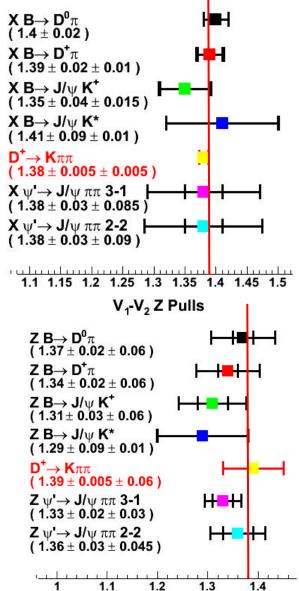


Scale Factor from V1-V2

- •Fit two independent subsets of 'primary' [I.e. non-B] tracks
- •Measure (x_1, y_1, z_1) and (x_2, y_2, z_2)
- •Obtain Δ/σ for x, y and z
- •Fit core with single gaussian (central value)
- •Repeat fit with two gaussians ('syst.')
- •Still using 1.38

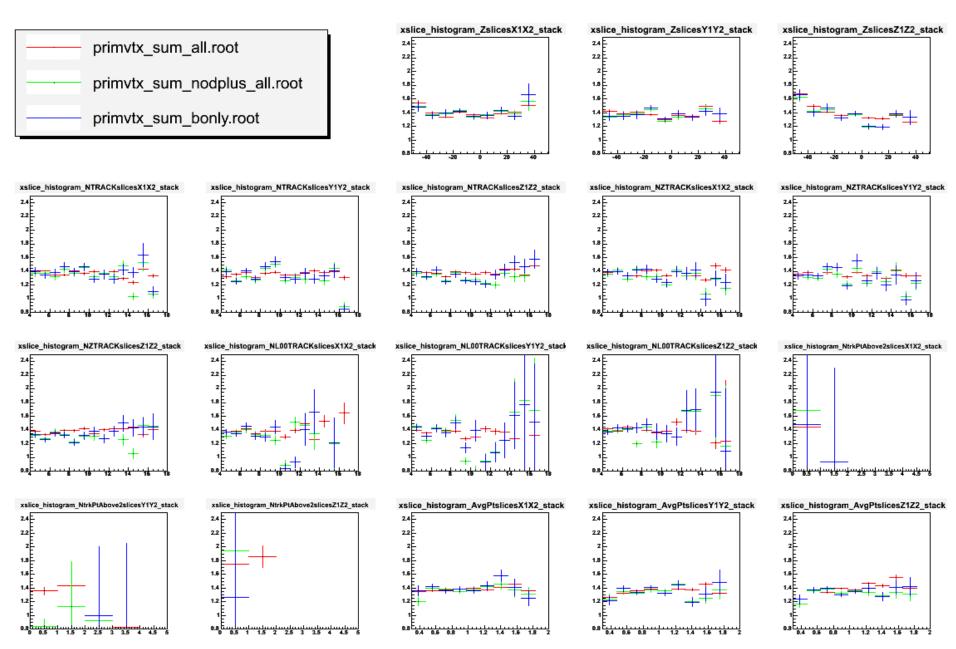
For what follows





V₁-V₂ X Pulls

Is the PVSF 'universal'?

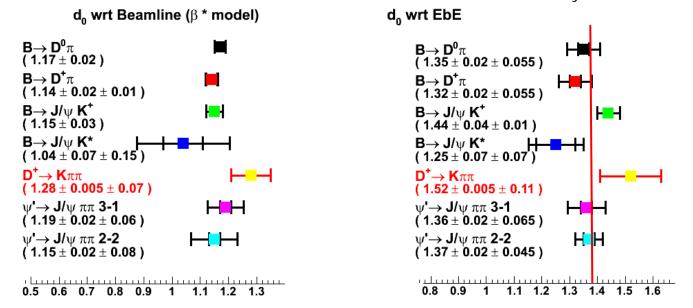


Bottomline...

- Scale factor shows no strong dependence on variables probed
- •Any other variable you want to see?
- •We could be more accurate but remember that the statistics is limited!
- Focus on assessing systematics
 - Inter-sample
 - Vary fit model (just like the overall fit)
- •Will try to improve a little more on statistics ($K\pi\pi\pi$, $D\pi\pi\pi$), but mostly aiming at improving systematics

Impact Parameters

•We can use the B I.P. pulls as cross check of the L_{xv} resolution...

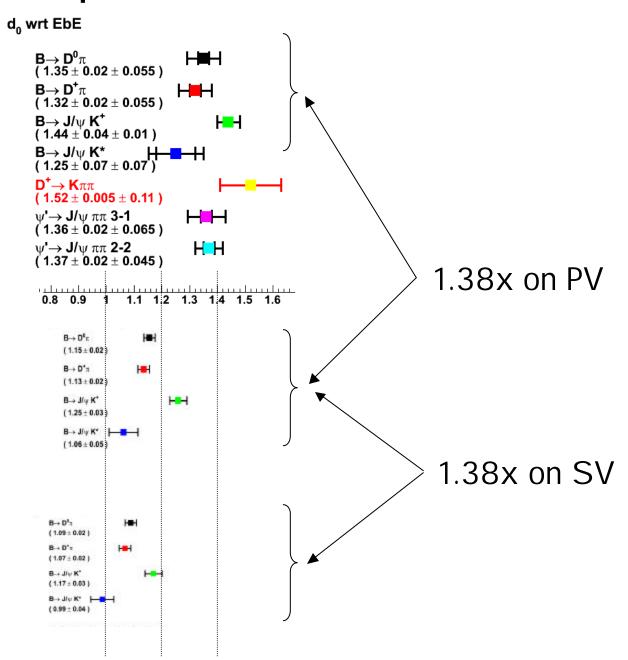


•Apply 1.38x and beamline constraint, check what happens:

For instance with $D^0\pi$: EbE: 1.35 \rightarrow 1.15 EbE+SSvtx: 1.26 \rightarrow 1.09

- •PV scale factor is not the full story: when you bring down σ_{PV} , the importance of σ_{SV} increases
- •Need to get the SV scale factor right!
- REM: even if PVSF=SVSF, we cannot use one common LxySF

Impact Parameters & scale factors

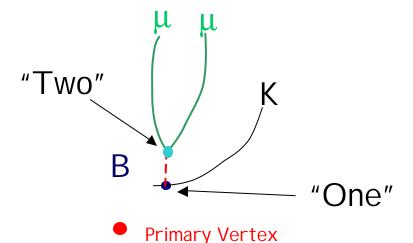


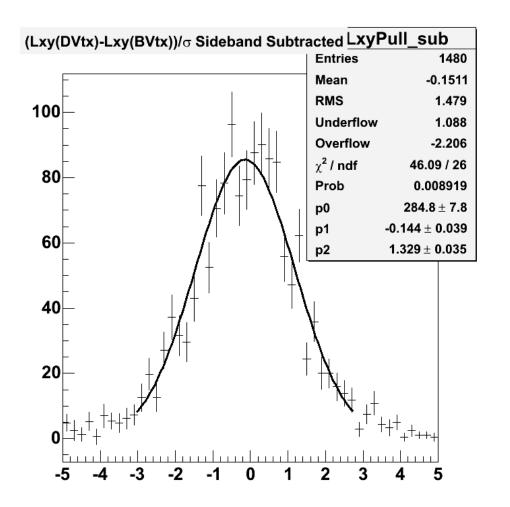
Secondary Vertex

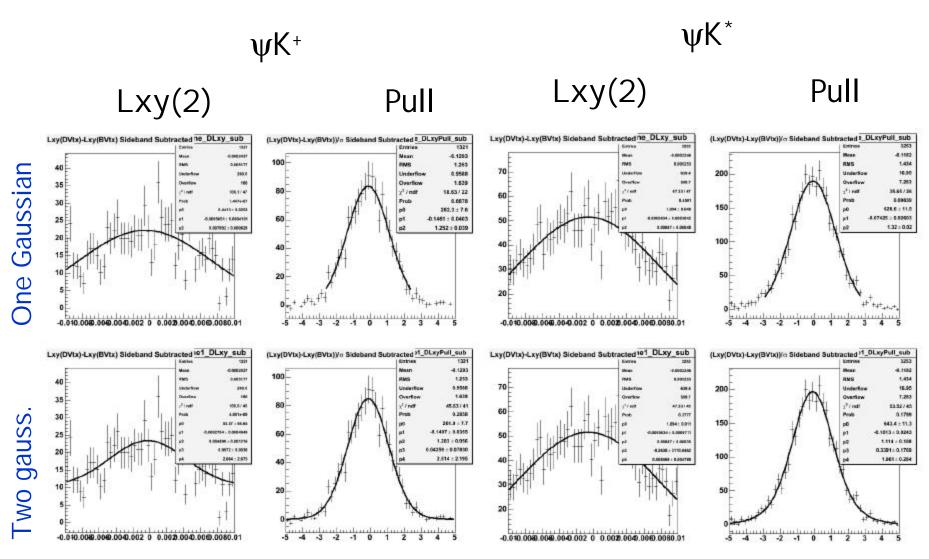
Scale factor from B decays

Example: B→ψK+

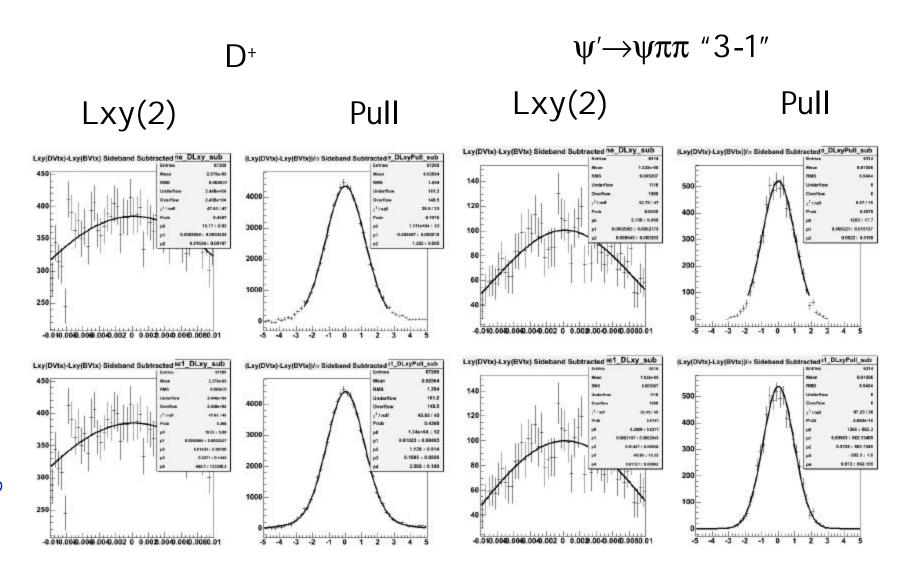
- •Fit ψ to a single vertex
- "point" ψ back to K
- •Measure L_{xy} wrt B vertex
- Pull is a proxy for a "seconday vertex" pull!





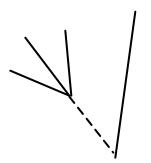


Charm...

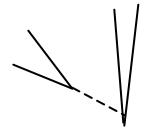


 ψ' can be used in two different ways to probe SV ψ'→ψππ "2-2"

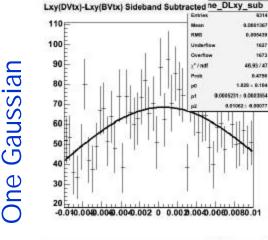
"3-1"



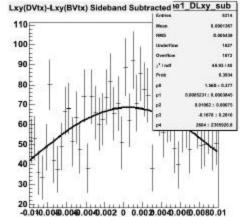
"2-2"



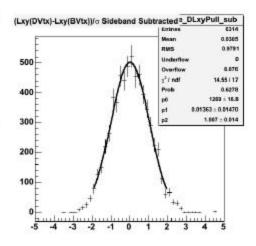
Two gauss

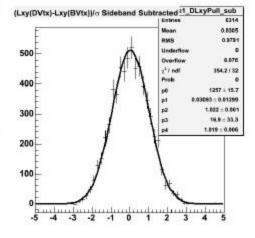


Lxy(2)

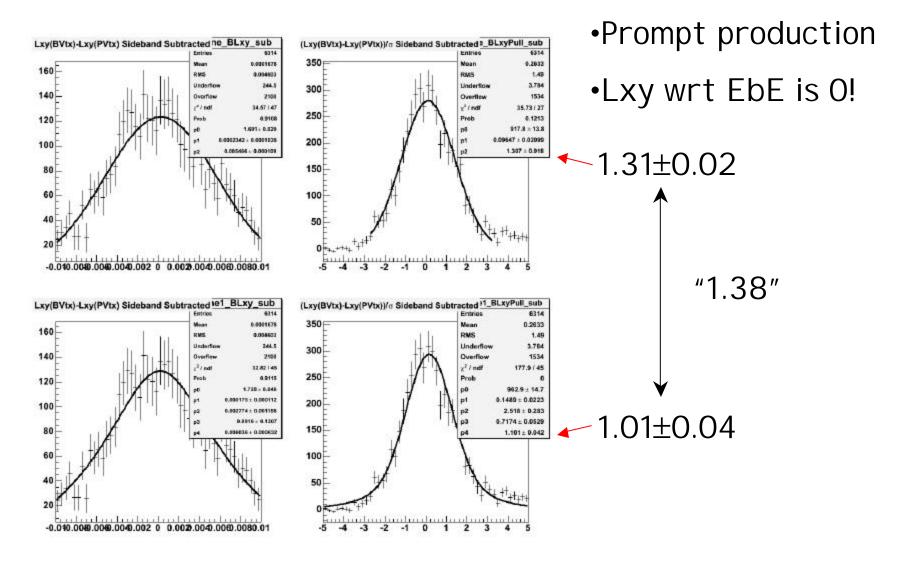




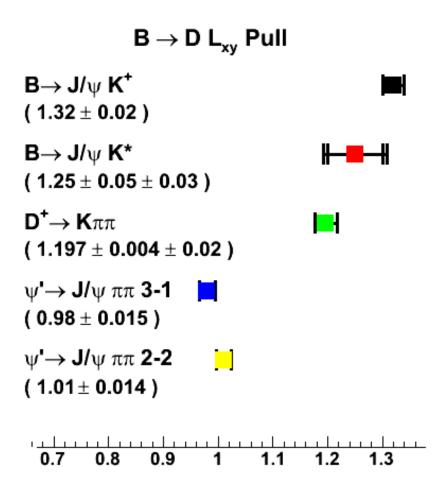




..and can be used also to probe the PV scale factor



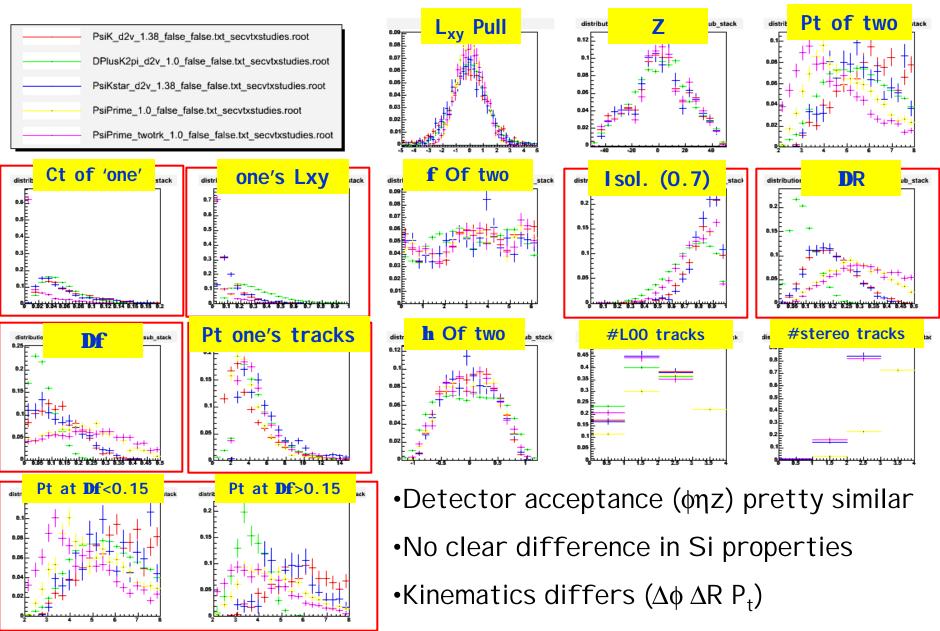
Bottomline for SV



- Hidden dependencies!
 - Detector acceptance?
 - Kinematics?
 - Multiplicity? (no: ψK*)
- Figure out which distributions are different
- 2. Check dependency!

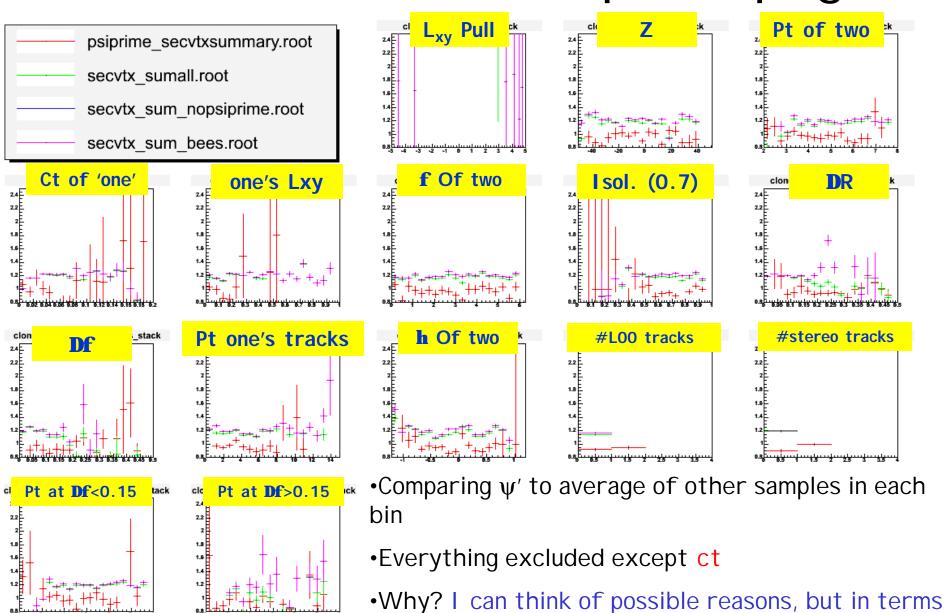
...results in the next pages

Distributions



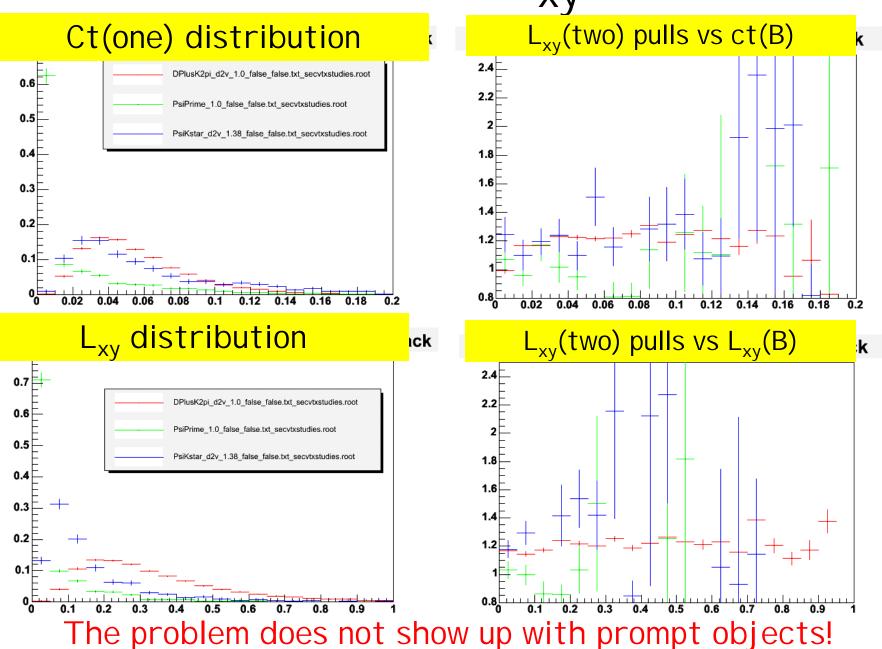
However...

Pulls vs variables in prev. page

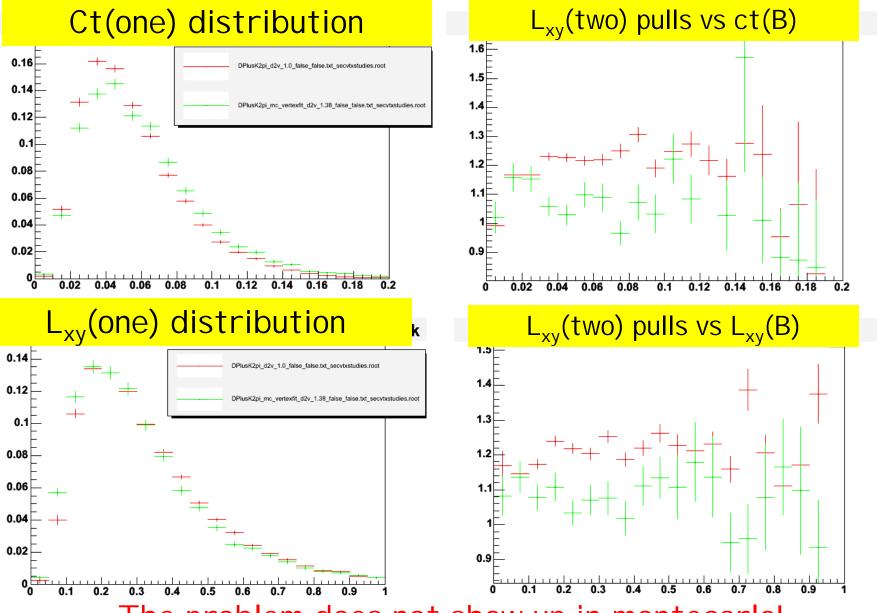


of bugs mostly! WORK IN PROGRESS

Ct and L_{xy}

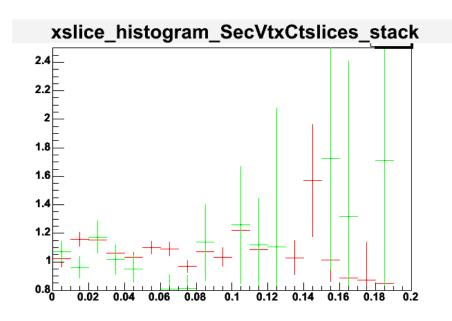


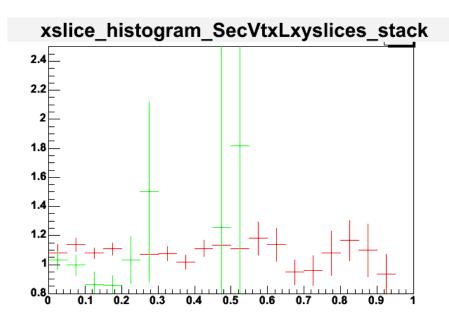
D+ Montecarlo vs data



The problem does not show up in montecarlo!

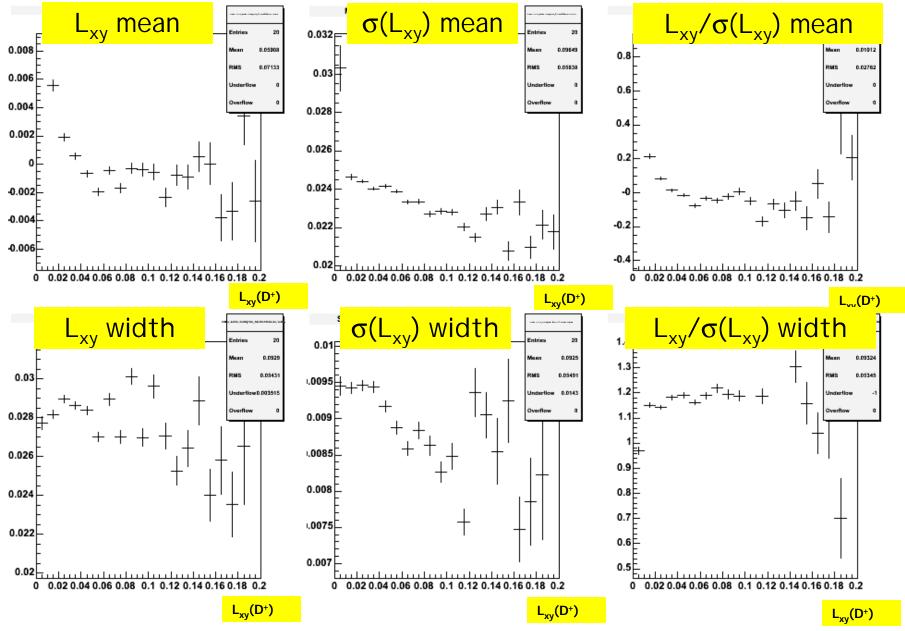
ψ' data vs D+ MC



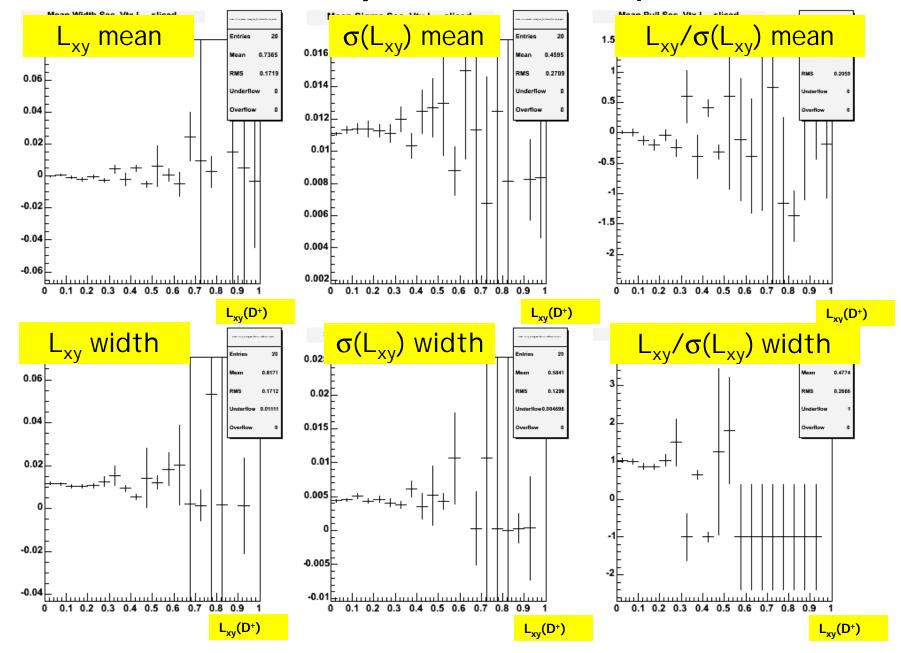


- They are compatible!
- The 'bug' affects non-prompt data only!!!

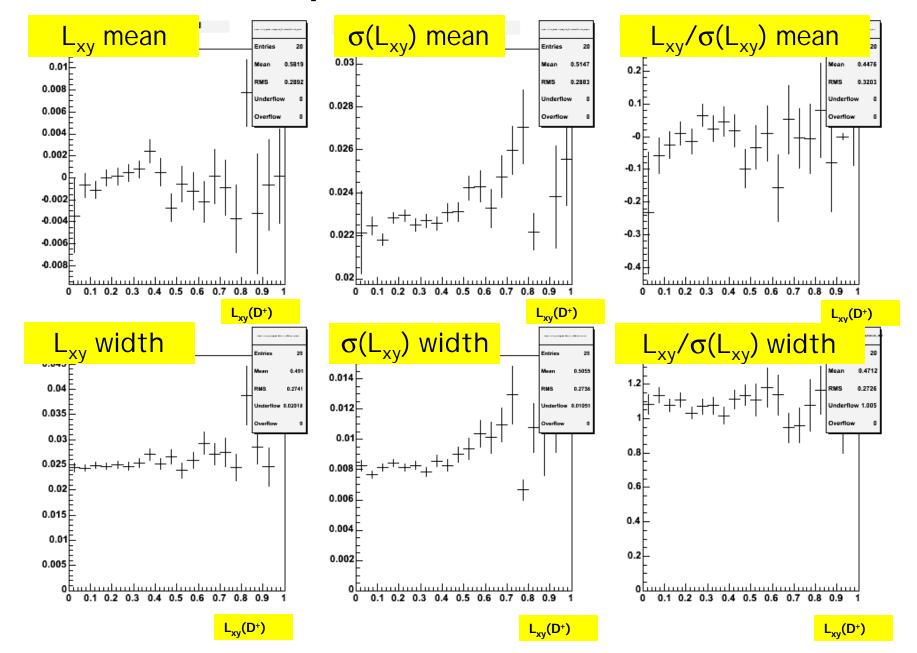
Is it in $\sigma(L_{xy})$ or in L_{xy} ?



Same plots for ψ'



Same plots for MC D⁺



Secondary Vertex Conclusions, so far

- We have enough statistics to study the SV scale factor
- Dependency on ct unexpected
 - Problem shows up 'only' in long lived signal in data
 - it's a pity that's what we want to use for our analyses;)
 - Semileptonic lifetime?
- Working on finding the cause
 - Montecarlo: It works!
 - Investigate other variables in data (Impact parameter?)
 - Probe other samples (D⁰?)
- Possible candidates:
 - Alignment
 - Scale factors?
- This is the last pending big issue at the moment

Moving along the plans for improvements!

- 1. Understand beamline parameterization:
 - I. Is it modeled correctly
 - II. Is it measured correctly
 - ⇒ Include our best knowledge of it!
- 2. Are secondary vertex pulls ok?
 - Check with montecarlo truth
 - II. Use n-prong vertices (J/ ψ K, K $\pi\pi^{+/0}$, K $\pi\pi\pi^{+/0}$)
- Investigate dependencies (Pt, z,multiplicity, η) with full statistics

Plan

- PV shows a very consistent picture
- Finish up PV studies (~days):
 - re-run some ntuples to get full statistics in all cases
 - Including Κπππ, Dπππ
 - Use also $B \rightarrow \psi K$ background to get a source of studies for prompt L_{xv} pulls?
- SV riddle to be solved!
- I am working full steam on this.
- Two more weeks according to schedule, to straighten everything out, document and insert in the blessing pipeline.

Backup

Outline

- Current status
 - What was used for the mixing results
 - What is the current understanding of Ebe
- Plans for improvements
 - How can we improve?

Current status

EbE: itearative track selection/pruning algorithm to provide an unbiased estimate of the PV position on an Event-by-Event basis

Hadronic analyses used a flat ~25um beamline!

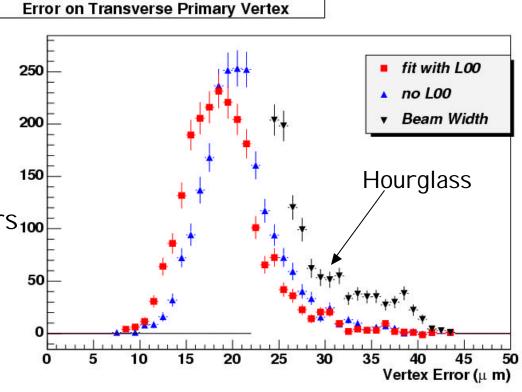
Possible improvements:

- Move to "hourglass"

Move to EbE

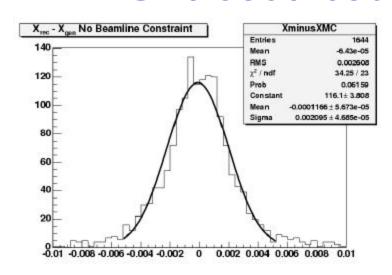
EbE + Hourglass

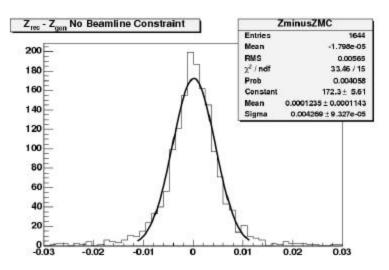
One of the ½ leptonic analyses used this with fixed hourglass parameters₁₀₀



What do we know about EbE?

Unbiased estimator of PVTX





Reasonable (~5%) control of systematics

Mode	x scale	y scale	z scale	
$B^{\pm} \rightarrow \psi K^{\pm}$	1.327 ± 0.035	1.399 ± 0.035	1.375 ± 0.029	
$B^{\pm} ightarrow D^0 \pi^{\pm}$	1.408 ± 0.030	1.398 ± 0.031	1.367 ± 0.29	
$B^0 o D^{\pm}\pi^{\mp}$	1.426 ± 0.034	1.336 ± 0.029	1.288 ± 0.027	

	Transverse	Z
Data (V ₁ -V ₂)	1.33±0.035	1.37±0.035
MC (V ₁ -V ₂)	1.192±0.034	1.26±0.035
MC (V-truth)	1.24±0.036	1.23±0.032
J/y Prompt Peak	1.236±0.024	~ND~
J/y d ₀ /s	1.176±0.019	~ND~

Cross checks using I.P.(B)

Pull on Impact Parameter

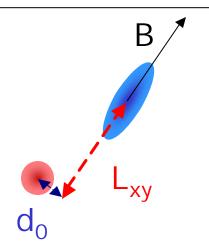
Mode	Beamline	Beamline	Event-by-Event	Event-by-Event	
	$\sigma = 25\mu$	z dependent σ	w/beam constraint	w/o beam constraint	
$B^{\pm} \rightarrow D^0 \pi^{\pm}$	1.297 ± 0.025	1.178 ± 0.039	1.202 ± 0.021	1.050 ± 0.025	
$B^0 \to D^{\pm} \pi^{\mp}$	1.256 ± 0.026	1.118 ± 0.027	1.163 ± 0.020	1.046 ± 0.027	

Z dep. Beamline improves pulls!

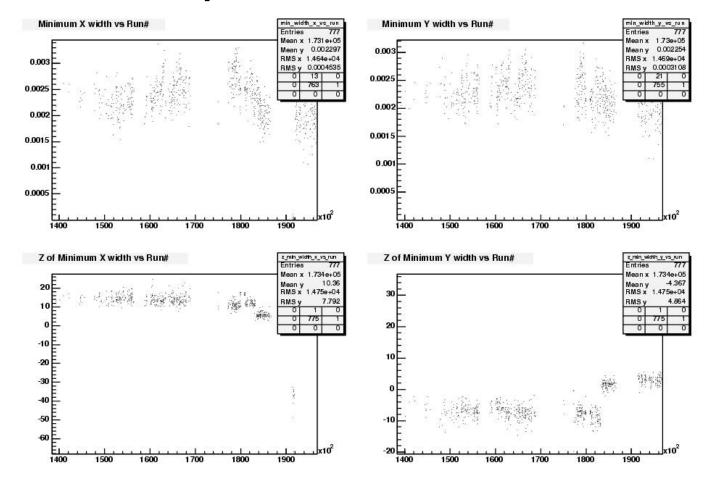
Something funny when beamline is used!

Scale factors work!

- L_{xy} involves three ingredients:
 - EbE
 - Secondary vertex
 - Beamline (in beamline constrained fits)



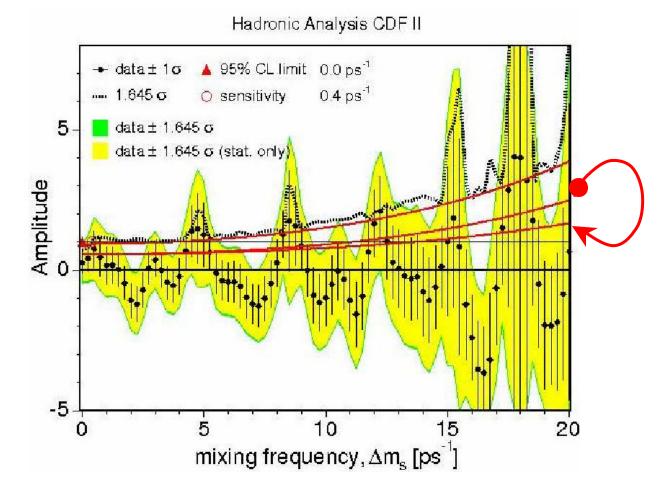
Time dependence of Hourglass parameters



Implementing DB access of time-dependent parameters

What do we gain?

- 1. 15-20% In vertex resolution!
- 2. Better control of systematics (hard to evaluate)
- 3. Correct EbE resolution (it is not clear that it is correct now)



•Red arrow is the effect of 1. Only

Euphemism

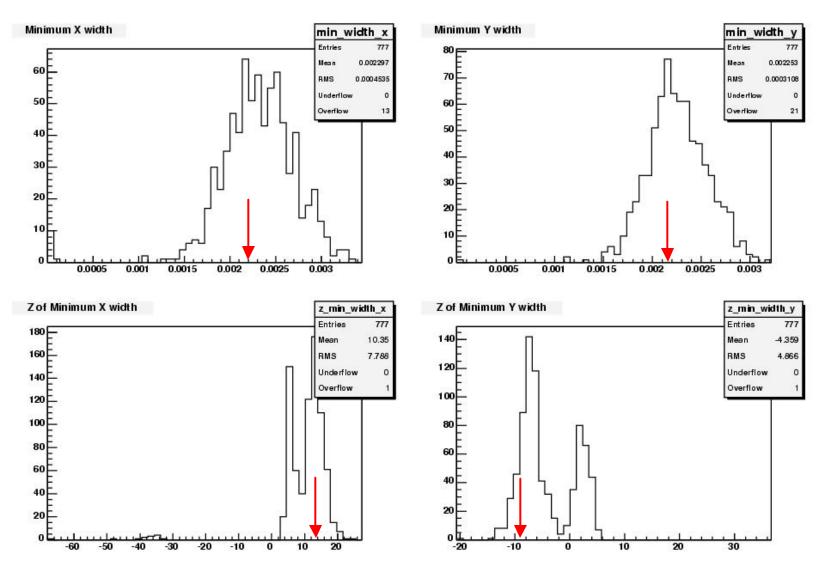
- •Point 2. Affects mostly the green area (tiny?)
- •Point 3. Has an effect qualitatively similar to 1., but hard to evaluate

Hadronic analysis systematics

source	selected Δm_s scan points				
	0.0	5.0	10.0	15.0	20.0
$B_s \to D_s K$ level	0.019	0.024	0.030	0.037	0.047
dilution scale factors	0.143	0.168	0.205	0.254	0.314
dilution templates	0.119	0.147	0.178	0.211	0.246
fraction of Λ_b	0.014	0.009	0.009	0.011	0.012
Punzi term for σ_{ct}	0.009	0.008	0.022	0.033	0.030
dilution of $B \to DX$	0.025	0.001	0.000	0.000	0.001
σ _{ct} scale factor	0.000	0.024	0.061	0.090	0.144
usage of L00 in bias curve	0.001	0.001	0.001	0.001	0.001
Bs lifetime uncertainty	0.001	0.001	0.001	0.001	0.001
reweighted p_t spectrum	0.001	0.001	0.001	0.001	0.001
non-Gaussian tails in ct resol.	0.001	0.027	0.052	0.078	0.104
neglect B^0 in fit	0.039	0.036	0.033	0.031	0.028
effect of $\Delta\Gamma/\Gamma = 0.2$	0.028	0.028	0.028	0.028	0.028
Total systematic	0.195	0.232	0.289	0.357	0.443
Statistical	0.393	1.129	1.010	2.652	5.281

Hourglass parameters from DB

Profiles



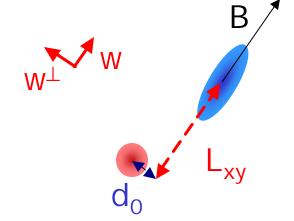
Relative PV/BV contribution to d_0 and L_{xv} pulls

$$\mathbf{S}_{L_{xy}}^{2} = {}^{t}w\mathbf{S}_{PV}^{2}w + {}^{t}w\mathbf{S}_{SV}^{2}w$$

$$\mathbf{S}_{d_{0}}^{2} = {}^{t}w^{\perp}\mathbf{S}_{PV}^{2}w^{\perp} + {}^{t}w^{\perp}\mathbf{S}_{SV}^{2}w^{\perp}$$

$$w = (x, y)$$

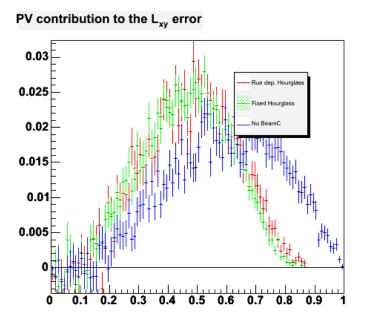
$$w^{\perp} = (y, -x)$$

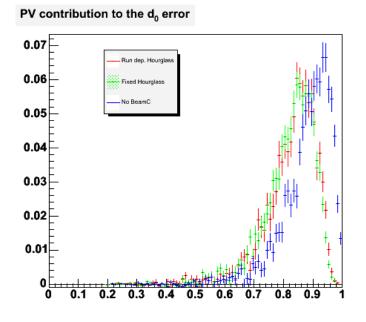


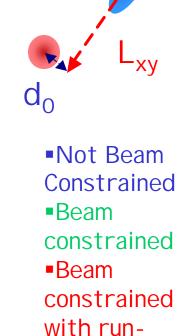
- •PV and BV are linear combinations of the same covariances (σ_{PV} , σ_{SV}), with different coefficients
- $^{\bullet}L_{xy}$ sensitive to the major axis of σ_{SV}
- •Relative weight of PV and SV covariances different for L_{xy} and d_0
- •Look at:

Note: the two L_{xy} (or d_0) pieces do not linearly add to 1!

Relative PV/BV contribution to IP and Lxy pulls

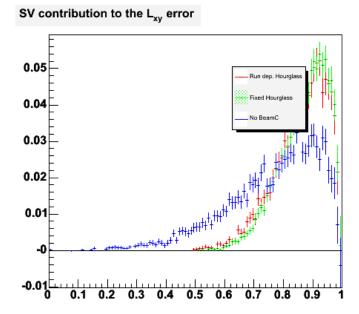


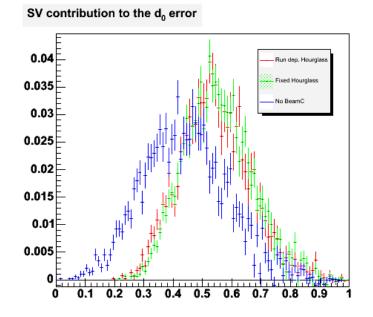




dep.

hourglass





Bottomline:

- SV and PV enter very differently in L_{xy} and d₀
- Relative contribution depends strongly on PV and SV scales
- Beam constraint squeezes the PV resolution significantly. Becomes second order on $L_{xy}!$
- We are in a regime where the SV scale factor is critical!

... now let's get more quantitative!